

INSTRUMENTAL EFFECTS

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Detector Saturation and Ringing, signal induced noise



How to detect?

Both effects can be detected through the presence of negative values above the cloud which are different from noise.

Necessary condition: background is correctly determined



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Negative values above the cloud



















Negative values above the cloud

CI51 @Lindenberg









CS135 @Lindenberg





TIME, UTC









Detection is possible in case the effect is large enough (sufficiently large range with negative values)

Correction?

Lufft wanted to investigate the effect in September and hopefully provide first (not final) info during SWG Ceilinex at Munich in October key: determination of dead time Vaisala wanted to investigate this effect we will see what will be the outcome Campell has undershooting, too be careful with device settings!

Small effects?

good question

Cloud calibration method

excludes clouds which produces such negative values





CLOUD BASE HEIGHT

-> Presentation by Uli Goersdorf





Not so low clouds, CBH ~ 1.9 km CEILINEX 2015, LINDENBERG, 20 Jul 2015 4.4 UTC ALL CHM140101 CHX080082 CS1 CL31RUB CL51CG 237 23 23 23 23 337 ATT BSC ATT. BSC ATT.88C CHM140101 CHM080882 ATT BSC ATT: 8SC - Zera Line - Zero Line - Zero Line - Zen Line - CBH - Marufacturer CBH - Metufacture EBH - Marufactur CBH - Marsdacture COH- Manufacturer - KH1 - CL31RUB - EL51CG 22 27 22 22 22 22 2.5 23 21 2.1 2.1 2.1 RANGE (ASL), km RANGE (ASL), Ium RANGE (ASL), km RANGE (ASL), km RANGE (ASL), km RANGE (ASL), km 1.9 tar 1.8 2.8 1.4-141 1.8 1.8 18 1.8 141 1.7 1.7 4 6 8 2 4 1 2 4 8 4 8 4 8 8 2 4 11 8 ÷. 18 0 .8 -2 . 6 0 -2 10 2. di la 6 ATT. BSC, a.u. 10-4 ATT. BSC, a.u. 10-4 ATT, BSC, a.u.=10-4 ATT: BSC, a.u.- 10-4 ATT, BSC, a.u.- 10-4 ATT. BSC, a.u. 10-4







low clouds, CBH ~ 0.5 km







NEW CBH SCHEME (1)









NEW CBH SCHEME (2)



Left: measured BSC profile together with all fit-tests of the artificial cloud shape Right: corresponding RMSD (the red dot corresponds to the beginning of the artificial cloud shape profile





Rain

Figures kindly provided by Margit









The first 2 or 3 or 4 bins cannot be used by any instrument, Neither Lufft ceilometer, neither Vaisala, neither Campbell

Example for Vaisala CL51 after start of dark current measurements





First few bins - Campbell









Problematic Lufft CHX-RAO device









Other Effects – tackled in other presentations

Dark current Measurements belly in free troposphere

Effect of Firmware changes

Effect of Parameter Settings for Vaisala and Campbell devices

- overlap (Vaisala)
- noise (Vaisala, Campbell)
- resolution (Campbell)

Overlap issue





CONCLUSIONS

Combined Effort of all is needed Close collaboration with manufacturers

For those who want to evaluate signals

- read the logbook
- have in mind all instrumental effects
- don't waste your time with tricky data
- => choose good data in the first place









